

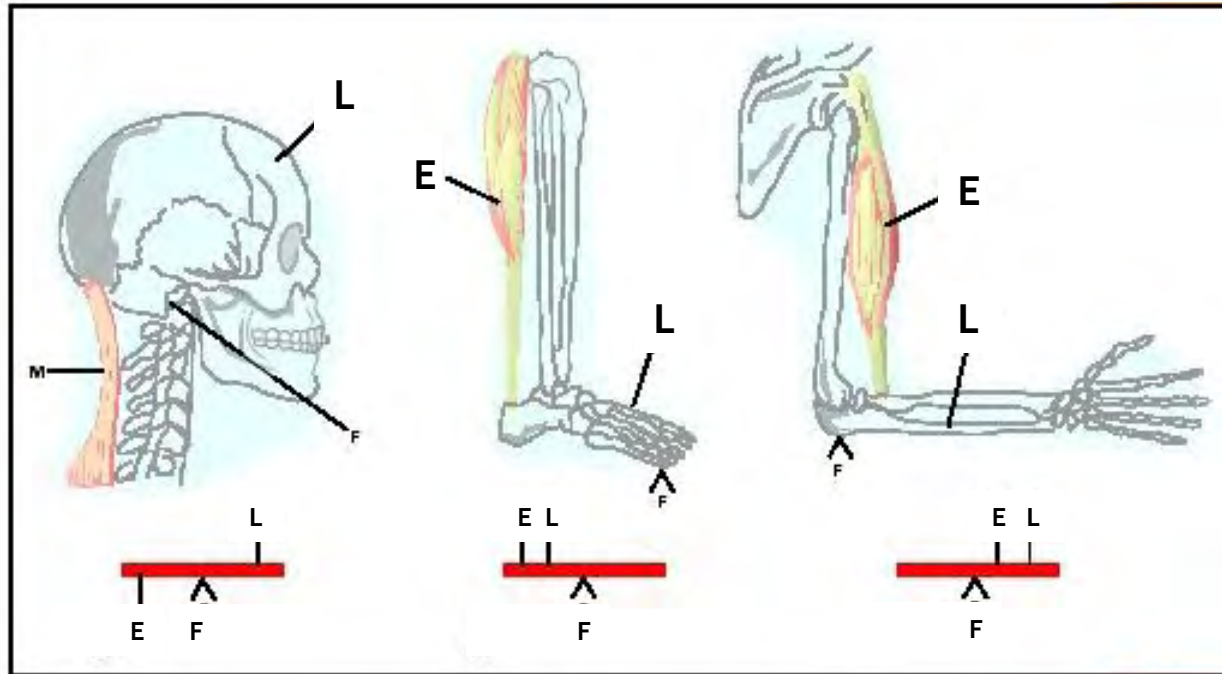
ROTATIONAL

EQUILIBRIUM

ROTATIONAL

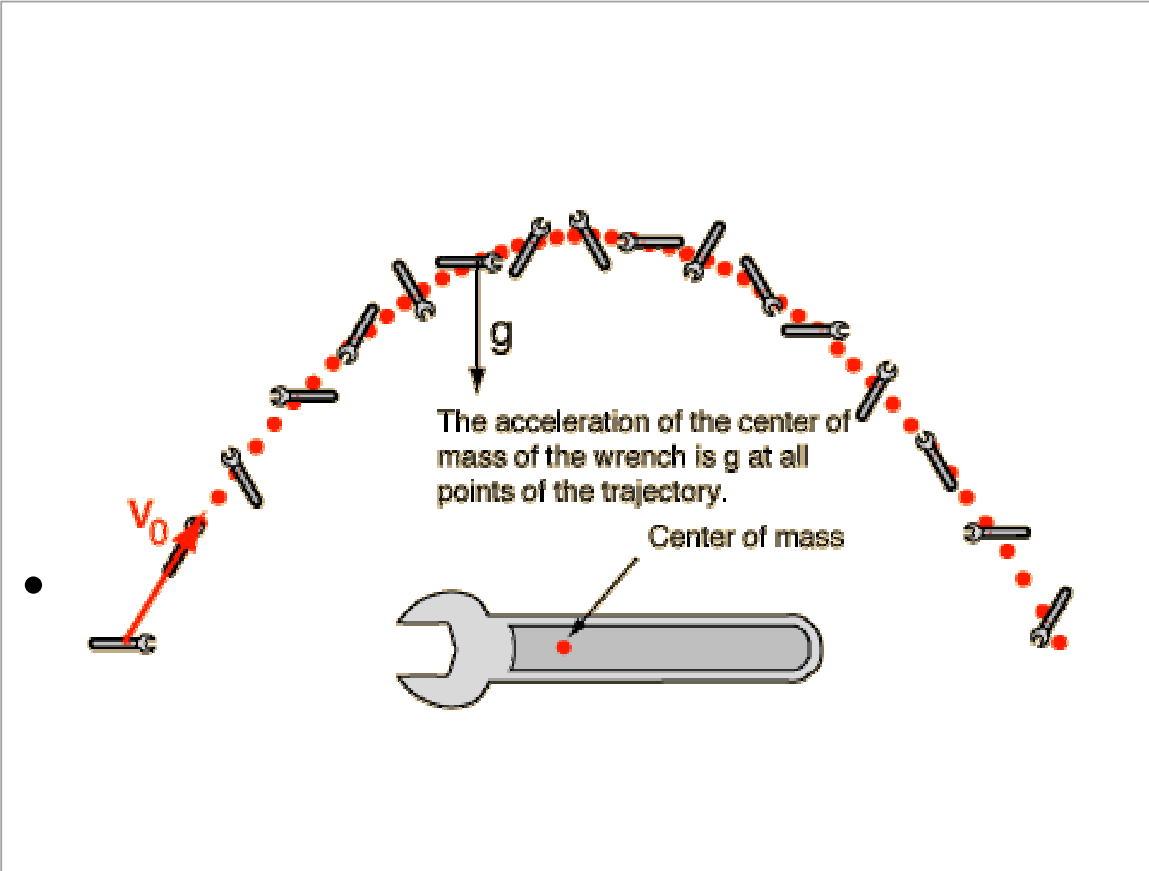
DYNAMICS





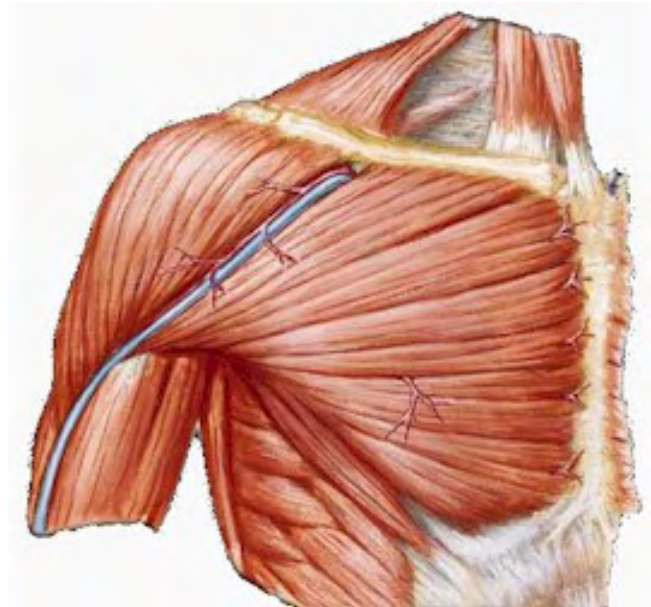
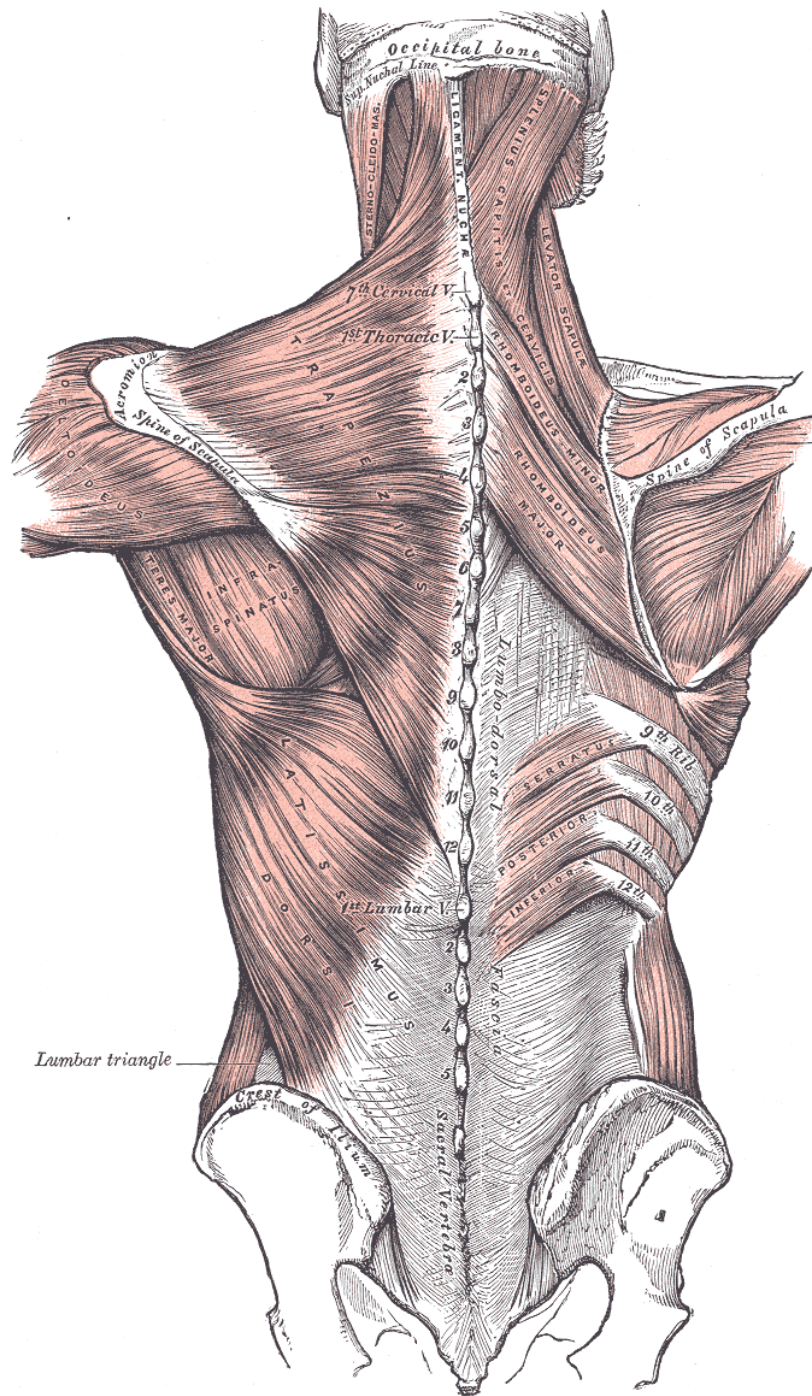
Classes of levers.

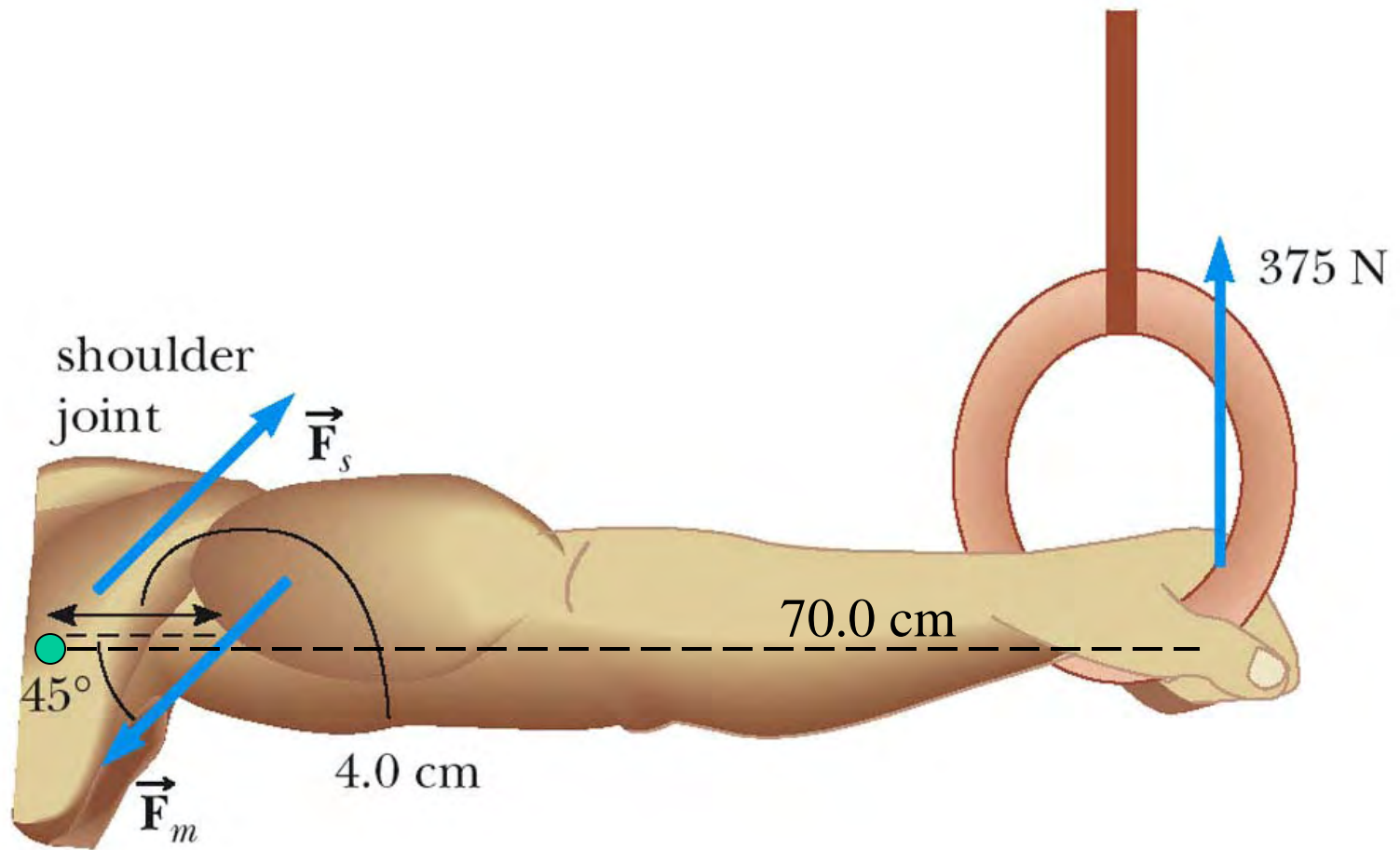
- (a) In a first-class lever, the fulcrum (F) is set up between the load (L) and the effort (E).
- (b) In a second-class lever, the resistance is between the fulcrum and the effort.
- (c) In a third-class lever, the effort is between the fulcrum and the load.











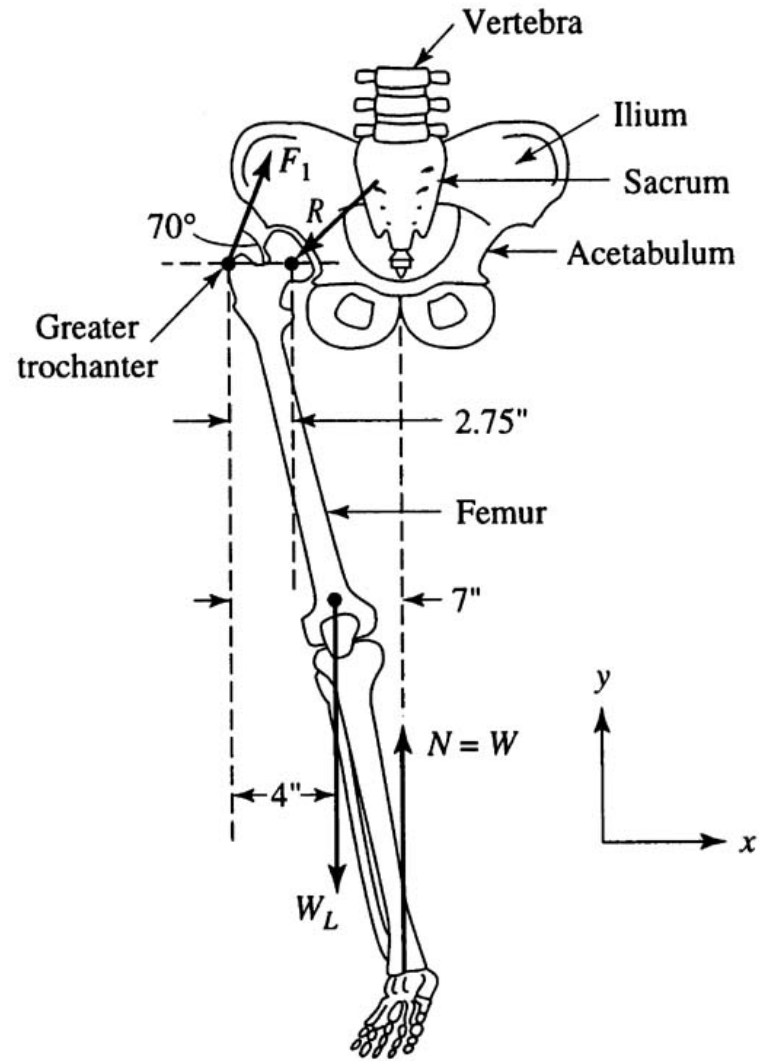


Figure 3.16. Anatomical sketch of leg and hip showing forces and dimensions. R is the force that the acetabulum exerts on the head of the femur. F_1 is the net force exerted by the hip abductor muscles. (Redrawn from Williams and Lissner, [4].)

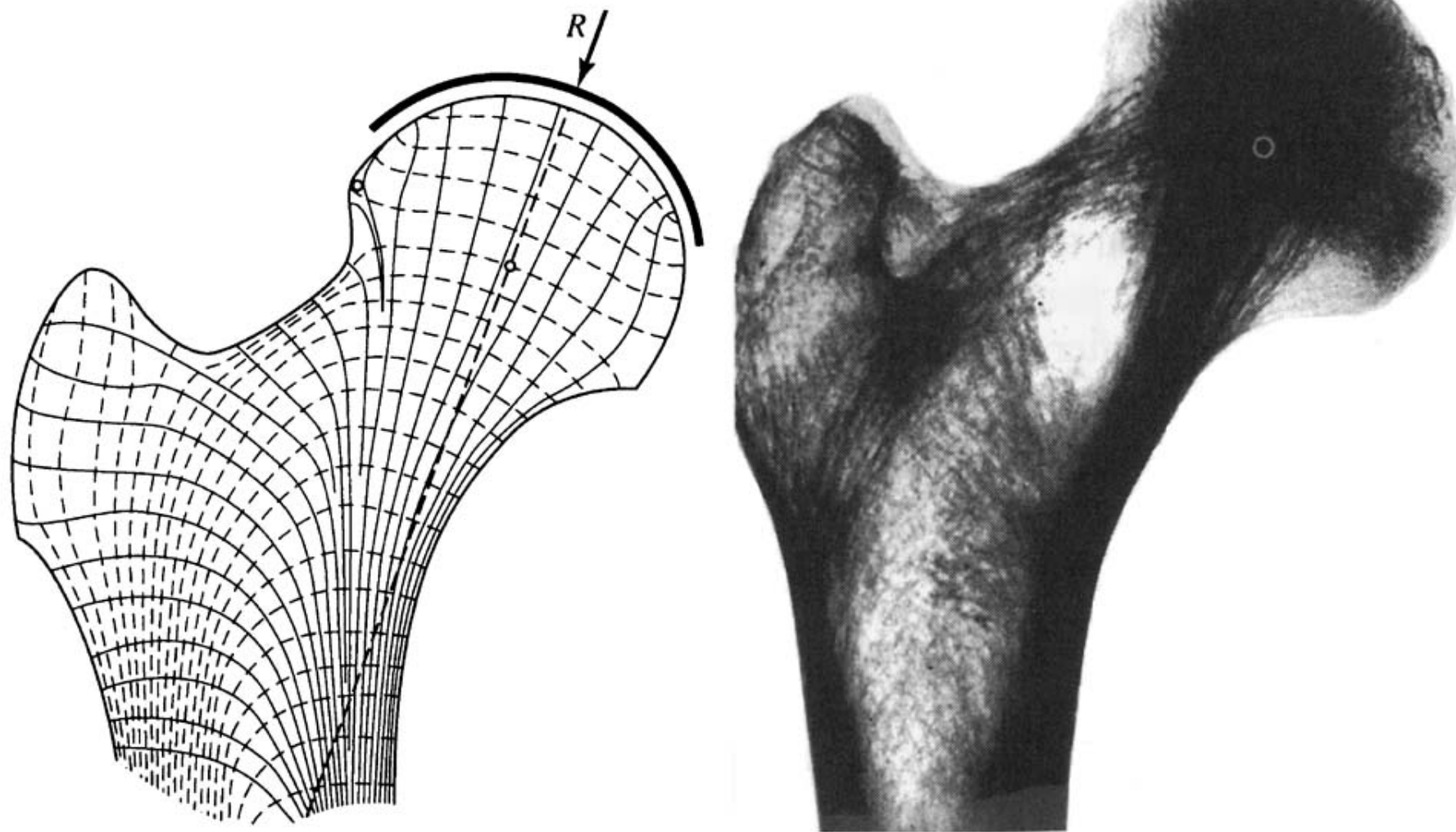


Figure 3.18. X-ray photograph of human femur.

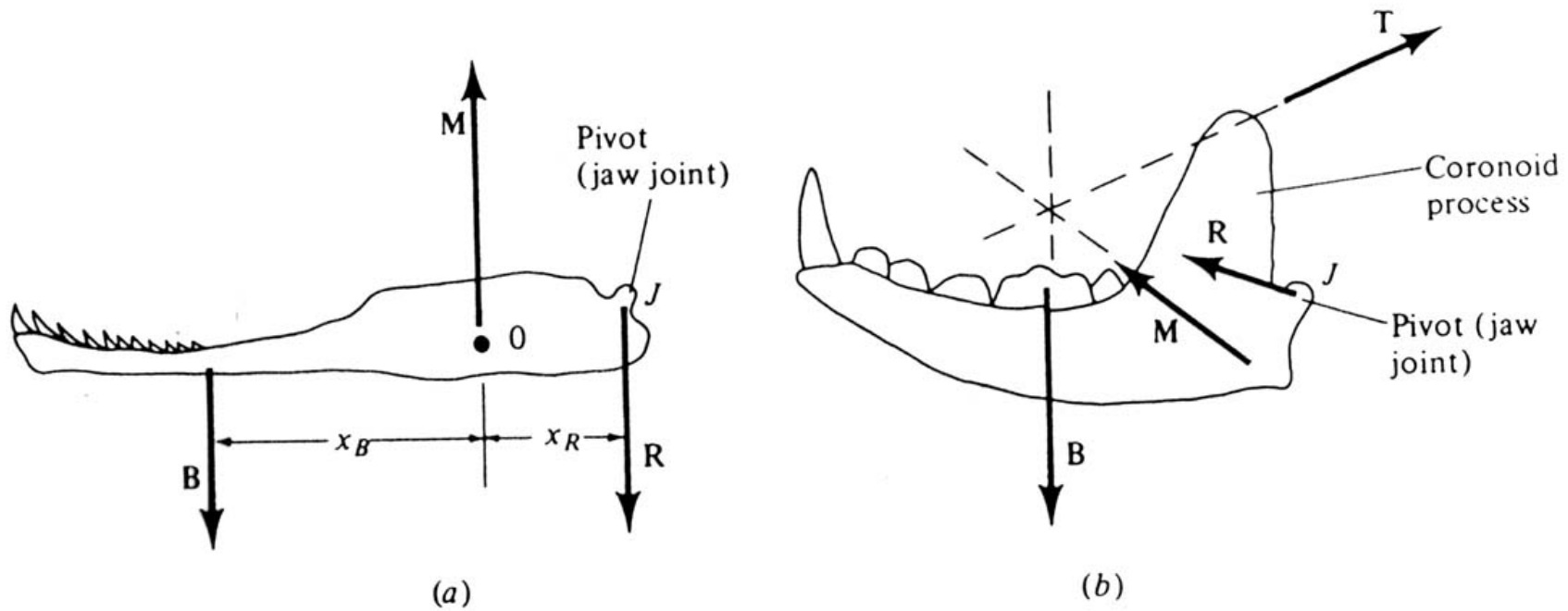
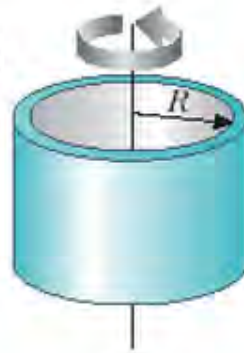


Figure 4.62. (a) Lower jaw of a primitive reptile. M is the force due to the muscle, B is the reaction force from the object being bitten, and R is the force due to the jaw joint at J . (b) A mammalian jaw. Muscle forces are shown as T and M . As explained in the text, the force R due to the jaw joint can be zero if the lines of action of the three forces T , B , and M intersect as shown here.

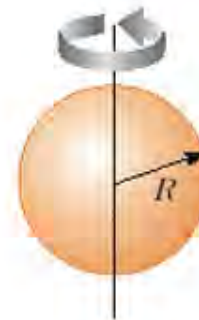
TABLE 8.1

**Moments of Inertia for Various Rigid Objects
of Uniform Composition**

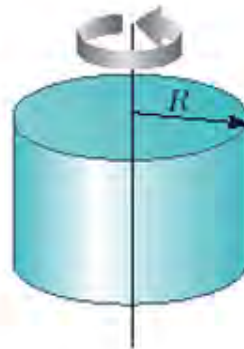
Hoop or thin
cylindrical shell
 $I = MR^2$



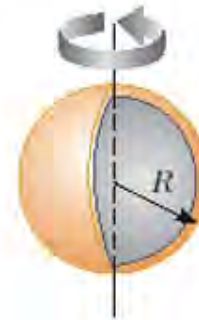
Solid sphere
 $I = \frac{2}{5} MR^2$



Solid cylinder
or disk
 $I = \frac{1}{2} MR^2$



Thin spherical
shell
 $I = \frac{2}{3} MR^2$

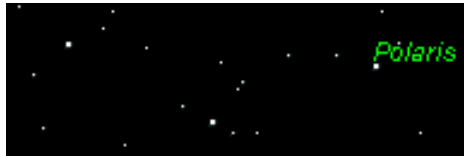


Long thin rod
with rotation axis
through center
 $I = \frac{1}{12} ML^2$

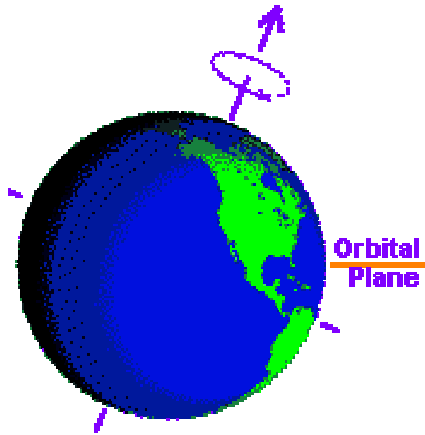


Long thin rod
with rotation axis
through end
 $I = \frac{1}{3} ML^2$



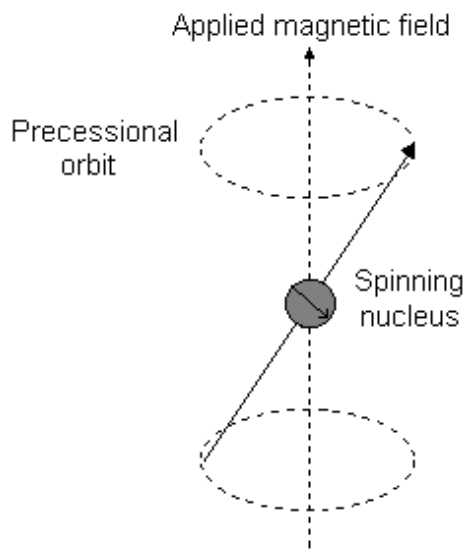


precession



The north pole doesn't always point at the North star (precession)



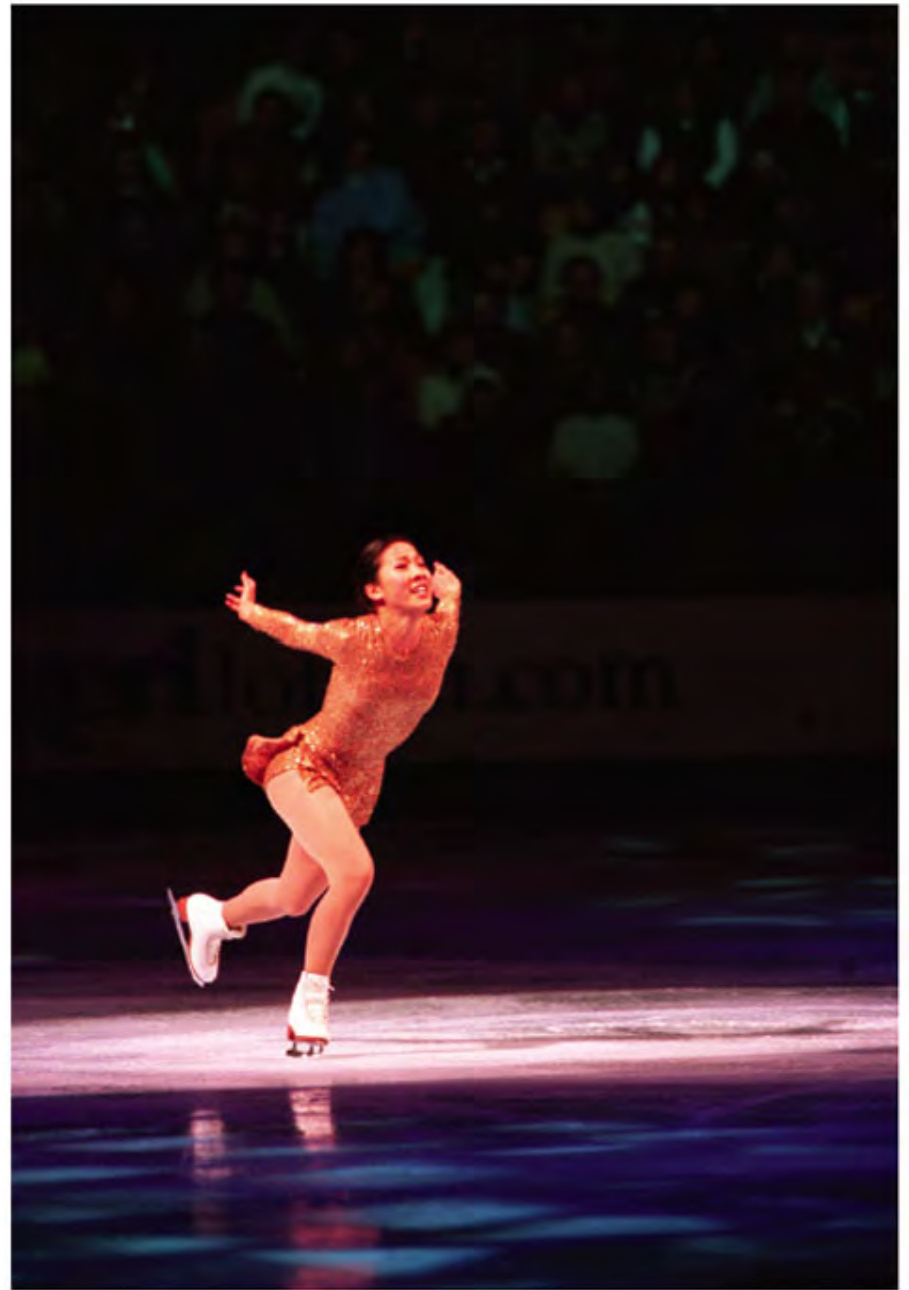


Precession of nuclei in a magnetic field = MRI (Magnetic Resonance Imaging)



Front legs tucked in, hind legs stretched out;
small I_{front} , big I_{back} ---- twist chest and head

Front legs stretched out, hind legs tucked in;
big I_{front} , small I_{back} ---- twist hindquarters



No. 6 on USA TODAY's 10 hardest things to do in sports is landing a quad in figure skating. Here's why it's hard, from a scientific perspective, from an expert athlete's perspective and from an average joe's.

By Gary Mihoces
USA TODAY



By Robert Hanashiro, USA TODAY

The quad toe loop is a delicate balance of height and rotation

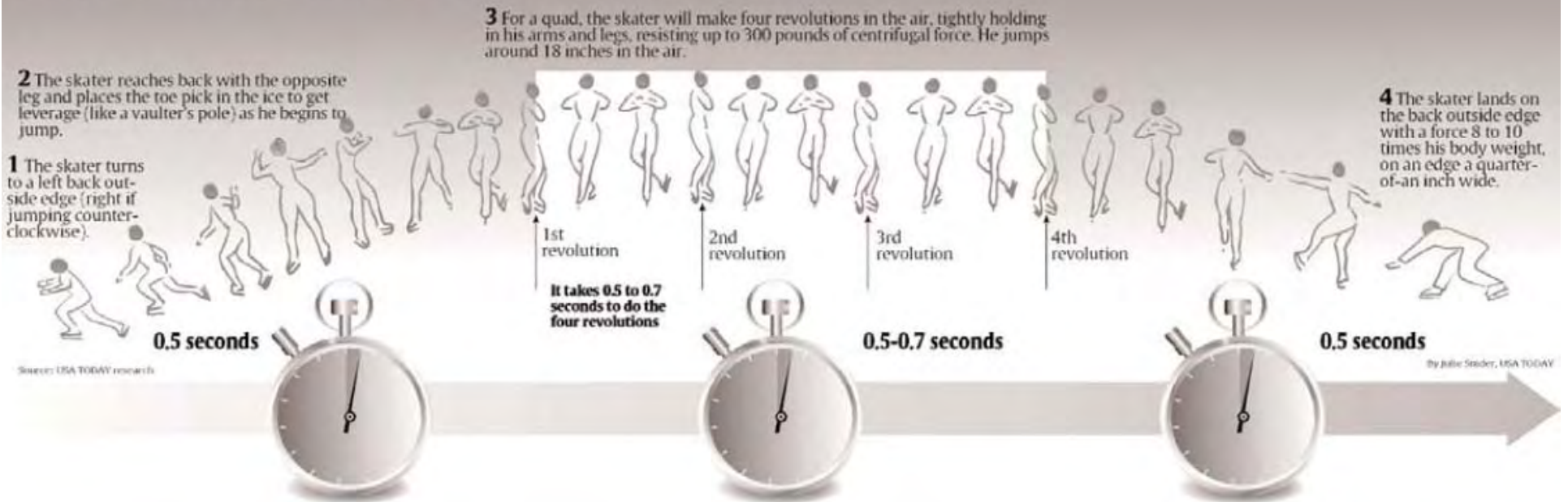
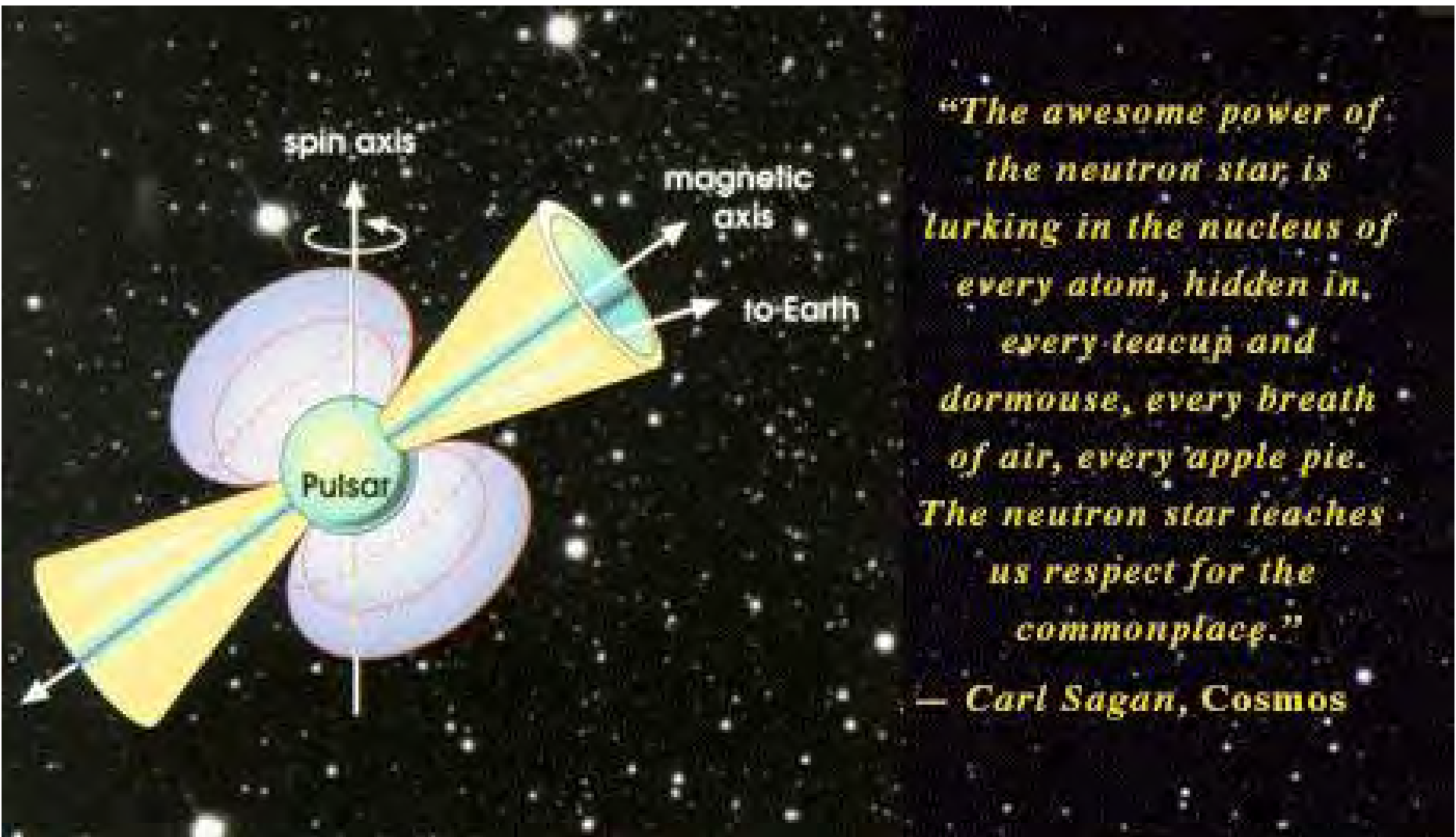


Table 3.1 Anthropometric Data

Segment	Definition	Segment Weight/ Total Body Weight	Center of Mass/ Segment Length		Radius of Gyration/ Segment Length			Density
			Proximal	Distal	C of G	Proximal	Distal	
Hand	Wrist axis/knuckle II middle finger	0.006 M	0.506	0.494 P	0.297	0.587	0.577 M	1.16
Forearm	Elbow axis/ulnar styloid	0.016 M	0.430	0.570 P	0.303	0.526	0.647 M	1.13
Upper arm	Glenohumeral axis/elbow axis	0.028 M	0.436	0.564 P	0.322	0.542	0.645 M	1.07
Forearm and hand	Elbow axis/ulnar styloid	0.022 M	0.682	0.318 P	0.468	0.827	0.565 P	1.14
Total arm	Glenohumeral joint/ulnar styloid	0.050 M	0.530	0.470 P	0.368	0.645	0.596 P	1.11
Foot	Lateral malleolus/head metatarsal II	0.0145 M	0.50	0.50 P	0.475	0.690	0.690 P	1.10
Leg	Femoral condyles/medial malleolus	0.0465 M	0.433	0.567 P	0.302	0.528	0.643 M	1.09
Thigh	Greater trochanter/femoral condyles	0.100 M	0.433	0.567 P	0.323	0.540	0.653 M	1.05
Foot and leg	Femoral condyles/medial malleolus	0.061 M	0.606	0.394 P	0.416	0.735	0.572 P	1.09
Total leg	Greater trochanter/medial malleolus	0.161 M	0.447	0.553 P	0.326	0.560	0.650 P	1.06
Head and neck	C7-T1 and 1st rib/ear canal	0.081 M	1.000	— PC	0.495	1.116	— PC	1.11
Shoulder mass	Sternoclavicular joint/ glenohumeral axis	—	0.712	0.288	—	—	—	1.04
Thorax	C7-T1/T12-L1 and diaphragm*	0.216 PC	0.82	0.18	—	—	—	0.92
Abdomen	T12-L1/L4-L5*	0.139 LC	0.44	0.56	—	—	—	—
Pelvis	L4-L5/greater trochanter*	0.142 LC	0.105	0.895	—	—	—	—
Thorax and abdomen	C7-T1/L4-L5*	0.355 LC	0.63	0.37	—	—	—	—
Abdomen and pelvis	T12-L1/greater trochanter*	0.281 PC	0.27	0.73	—	—	—	1.01
Trunk	Greater trochanter/ glenohumeral joint*	0.497 M	0.50	0.50	—	—	—	1.03
Trunk head neck	Greater trochanter/ glenohumeral joint*	0.578 MC	0.66	0.34 P	0.503	0.830	0.607 M	—
HAT	Greater trochanter/ glenohumeral joint*	0.678 MC	0.626	0.374 PC	0.496	0.798	0.621 PC	—
HAT	Greater trochanter/mid rib	0.678	1.142	—	0.903	1.456	—	—

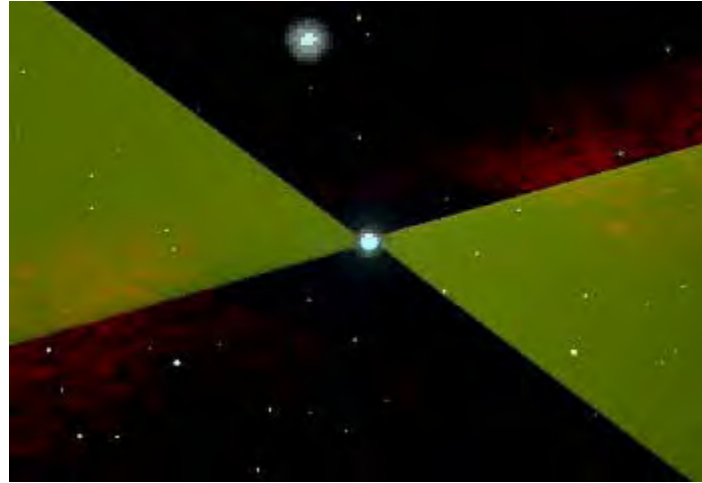
*NOTE: These segments are presented relative to the length between the greater trochanter and the glenohumeral joint.

Source Codes: M, Dempster via Miller and Nelson; *Biomechanics of Sport*, Lea and Febiger, Philadelphia, 1973. P, Dempster via Plagenhoef; *Patterns of Human Motion*, Prentice-Hall, Inc. Englewood Cliffs, N.J., 1971. L, Dempster via Plagenhoef from living subjects; *Patterns of Human Motion*, Prentice-Hall, Inc., Englewood Cliffs, N.J., 1971. C, Calculated.



“The awesome power of the neutron star, is lurking in the nucleus of every atom, hidden in, every teacup and dormouse, every breath of air, every apple pie. The neutron star teaches us respect for the commonplace.”

— Carl Sagan, Cosmos

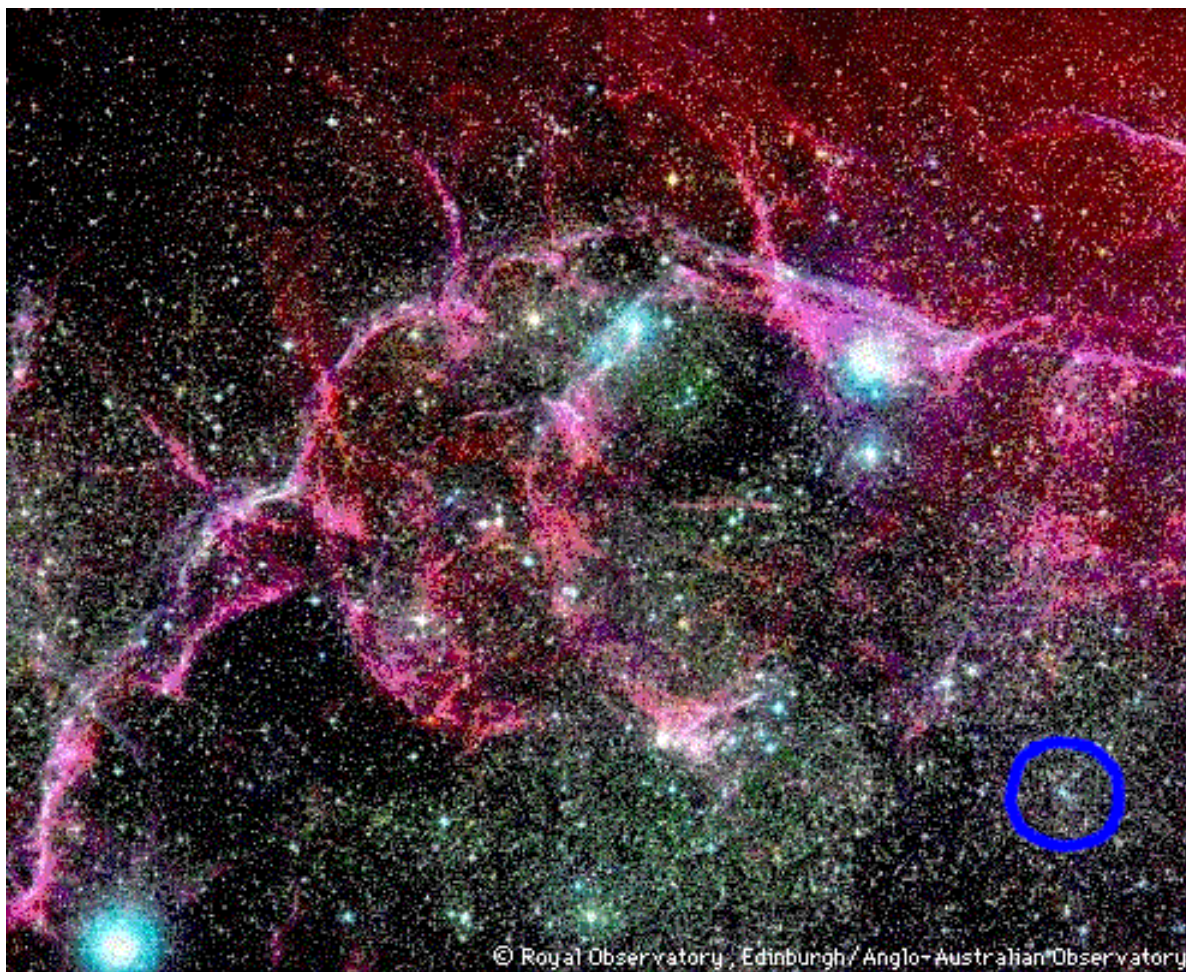


PSR B0329+54

This pulsar is a typical, normal pulsar, rotating with a period of 0.714519 seconds, i.e. close to 1.40 rotations/sec.



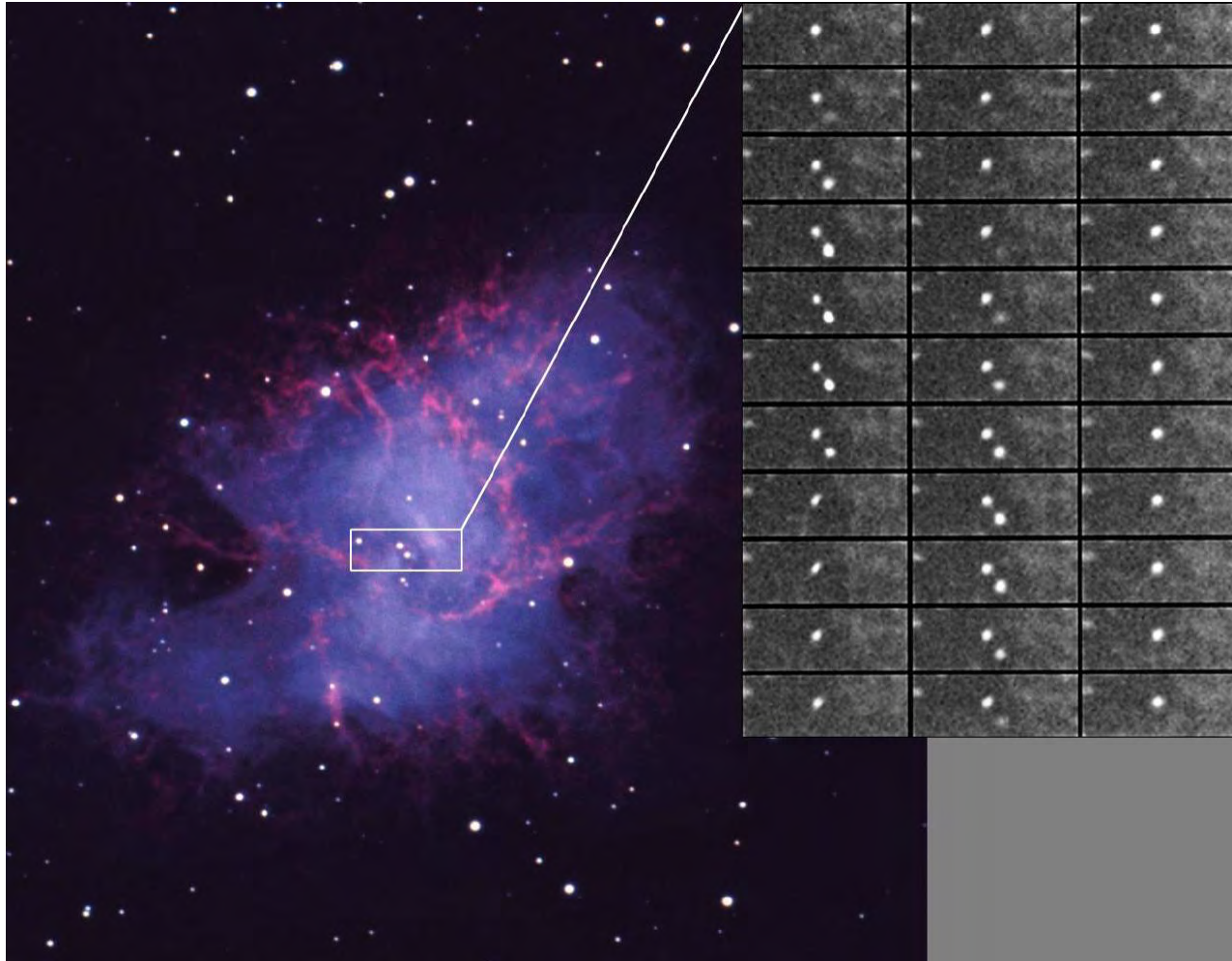
PSR B0833-45, The Vela Pulsar



This pulsar lies near the center of the Vela supernova remnant, which is the debris of the explosion of a massive star about 10,000 years ago. The pulsar is the collapsed core of this star, rotating with a period of 89 milliseconds or about 11 times a second.



PSR B0531+21, The Crab Pulsar

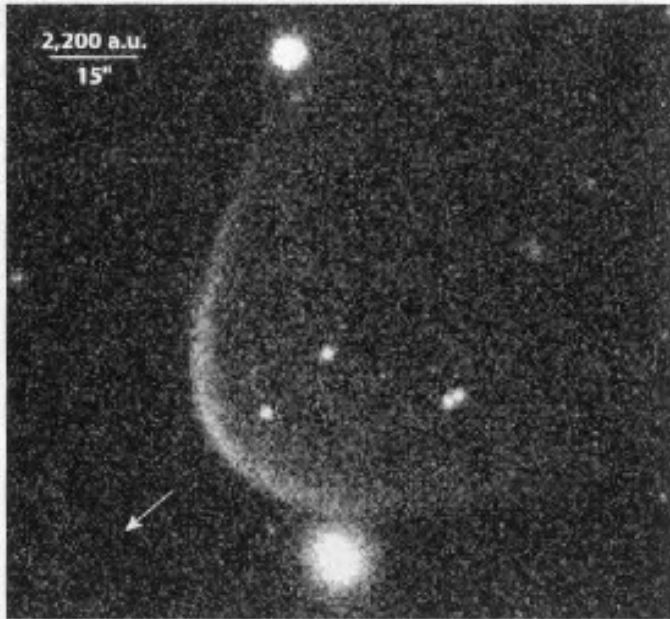


This is the youngest known pulsar and lies at the centre of the Crab Nebula, the supernova remnant of its birth explosion, which was witnessed by Europeans and Chinese in the year 1054 A.D. as a day-time light in the sky. The pulsar rotates about 30 times a second.



PSR J0437-4715

This is a recently discovered millisecond pulsar, an old pulsar which has been spun up by the accretion of material from a binary companion star as it expands in its red giant phase. The accretion process results in orbital angular momentum of the companion star being converted to rotational angular momentum of the neutron star, which is now rotating about 174 times a second.



PSR B1937+21

This is the fastest known pulsar, rotating with a period of 0.00155780644887275 seconds, or about 642 times a second. The surface of this star is moving at about 1/7 of the velocity of light and illustrates the enormous gravitational forces which prevent it flying apart due to the immense centrifugal forces.

